

Lecture 10: Supplemental Notes

Note Title

3/1/2006

Two calculations:

- (1) Definition of Γ^M_{np}
- (2) Relation of geodesic equation to the Euler-Lagrange Equations

1. Definition of Γ^M_{np}

$$\text{Given } \frac{\partial \xi^\alpha}{\partial x^\mu} = e^\alpha_\mu, \quad d\xi^\alpha = e^\alpha_\mu dx^\mu,$$

$$g_{\mu\nu} = \eta_{\alpha\beta} e^\alpha_\mu e^\beta_\nu, \quad + \quad \boxed{\partial_\mu e^\alpha_\nu = \partial_\nu e^\alpha_\mu}$$

This is symmetry of partial derivatives. i.e. $\frac{\partial}{\partial x} \frac{\partial}{\partial y} = \frac{\partial}{\partial y} \frac{\partial}{\partial x}$

Want to show

$$\Gamma^M_{np} = \frac{1}{2} g^{\mu\sigma} \{ \partial_\nu g_{\rho\sigma} + \partial_\rho g_{\nu\sigma} - \partial_\sigma g_{\mu\nu} \}$$

$$\partial_\rho g_{\nu\sigma} = \eta_{\alpha\beta} [(\partial_\rho e^\alpha_\nu) e^\beta_\sigma + e^\alpha_\nu \partial_\rho e^\beta_\sigma]$$

$$\partial_\nu g_{\rho\sigma} = \eta_{\alpha\beta} [(\partial_\nu e^\alpha_\rho) e^\beta_\sigma + e^\alpha_\rho \partial_\nu e^\beta_\sigma]$$

$$\partial_\sigma g_{\mu\nu} = \eta_{\alpha\beta} [(\partial_\sigma e^\alpha_\mu) e^\beta_\nu + e^\alpha_\mu \partial_\sigma e^\beta_\nu]$$

$$\partial_z g_{\rho\sigma} + \partial_\rho g_{z\sigma} - \partial_\sigma g_{z\rho} = \mu_{\alpha\beta} \left[(\partial_\rho e_z^\alpha) e_\sigma^\beta + \cancel{e_z^\alpha \partial_\rho e_\sigma^\beta} + (\partial_z e_\rho^\alpha) e_\sigma^\beta + \cancel{e_\rho^\alpha \partial_z e_\sigma^\beta} - (\partial_\sigma e_z^\alpha) e_\rho^\beta - \cancel{e_z^\alpha \partial_\sigma e_\rho^\beta} \right]$$

$\alpha \Leftrightarrow \beta$ sym
plus red box
above

$$= 2 \mu_{\alpha\beta} (\partial_\rho e_z^\alpha) e_\sigma^\beta = 2 \mu_{\alpha\beta} e_\sigma^\beta \partial_\rho e_z^\alpha$$

Now $g_{\mu\sigma} = \mu_{\alpha\beta} e_\mu^\alpha e_\sigma^\beta$ implies

$$g_{\mu\sigma} e_\alpha^\mu = \mu_{\alpha\beta} e_\sigma^\beta, \text{ so}$$

$$2 \mu_{\alpha\beta} e_\sigma^\beta \partial_\rho e_z^\alpha = 2 g_{\mu\sigma} e_\alpha^\mu \partial_\rho e_z^\alpha$$

$$\partial_z g_{\rho\sigma} + \partial_\rho g_{z\sigma} - \partial_\sigma g_{z\rho} = 2 g_{\mu\sigma} e_\alpha^\mu \partial_\rho e_z^\alpha$$

$$\frac{1}{2} g^{\mu\sigma} \left\{ \partial_z g_{\rho\sigma} + \partial_\rho g_{z\sigma} - \partial_\sigma g_{z\rho} \right\} = e_\alpha^\mu \partial_\rho e_z^\alpha = \Gamma_{z\rho}^\mu$$

as we wanted to show.

2. Geodesics and Lagrangians

- From the coordinate transformation b/w a freely falling coordinate system ξ^M and a general coordinate x^a , we saw that a freely falling particle on a worldline $\xi^M(\tau)$ w/ $\ddot{\xi}^M(\tau) = 0$ + $g_{ab} \dot{\xi}^a \dot{\xi}^b > 0$ (timelike worldline), was observed to move according to

$$\ddot{x}^a + \Gamma_{bc}^a \dot{x}^b \dot{x}^c = 0 \quad - \text{the geodesic eqn.}$$

$$\text{w/ } \Gamma_{bc}^a = \frac{1}{2} g^{ad} \{ g_{bd,c} + g_{cd,b} - g_{bc,d} \}$$

- Here, we want to show that the paths given by the geodesic equation extremize the "length" of the path.

$$S = \int_{\tau_0}^{\tau_1} d\tau \, g_{ab} \dot{x}^a \dot{x}^b \Rightarrow \mathcal{L} = \frac{1}{2} g_{ab} \dot{x}^a \dot{x}^b$$

The Euler-Lagrange Eqn is

$$\frac{\partial \mathcal{L}}{\partial x^a} = \frac{d}{d\tau} \frac{\partial \mathcal{L}}{\partial \dot{x}^a}$$

$$\frac{1}{2} g_{ab,c} \dot{x}^a \dot{x}^b = \frac{d}{dt} (g_{ab} \dot{x}^a \delta_c^b) = \frac{d}{dt} (g_{ac} \dot{x}^a)$$

$$\frac{1}{2} g_{ab,c} \dot{x}^a \dot{x}^b = g_{ac} \ddot{x}^a + g_{ac,d} \dot{x}^a \dot{x}^d$$

Now $g_{ac,d} \dot{x}^a \dot{x}^c = \frac{1}{2} g_{ac,d} \dot{x}^a \dot{x}^d + \frac{1}{2} g_{dc,a} \dot{x}^d \dot{x}^a$

(just by "dummy" relabeling)

So

$$g_{ac} \ddot{x}^a + \frac{1}{2} (g_{ac,d} \dot{x}^a \dot{x}^d + g_{dc,a} \dot{x}^d \dot{x}^a - g_{ad,c} \dot{x}^a \dot{x}^c) = 0$$

Multiply by g^{ce}

$$g^{ce} g_{ac} = \delta_a^e \Rightarrow g^{ce} g_{ac} \dot{x}^a = \dot{x}^e$$

$$\ddot{x}^e + \frac{1}{2} g^{ec} \{ g_{ac,d} + g_{dc,a} - g_{ad,c} \} \dot{x}^a \dot{x}^d = 0$$

$$\boxed{\ddot{x}^e + \Gamma_{ad}^e \dot{x}^a \dot{x}^d = 0}$$

Important Points

So the application of the E-L equation to $\mathcal{L} = \frac{1}{2} g_{ab} \dot{x}^a \dot{x}^b$ yields the geodesic equation.

Since we have extremized the square metric interval using the E-L equation, this calculation formally proves that objects follow the extremal curve of space time as their natural motion.