

Nothing = $|000000\rangle$

Fock # space
Hilbert

BIG BANG

$\Delta E \Delta t \geq \hbar$

$\Delta t = \text{wage} \Rightarrow 's$
 $\Delta E = \text{small}$

$\sim 700,000$ years

Inflation

$3000^\circ K$
e⁻'s condense
onto p = H atom
so particle
density ↓
universe
transparent
CMBR

+ Energy + - E potential = 0
 $E_k + \text{Mass}$ gravitational

how does
particle + antipart
get into
this
energy
scheme?

OUR UNIVERSE
13.5 Billion years

nucleo-
synthesis

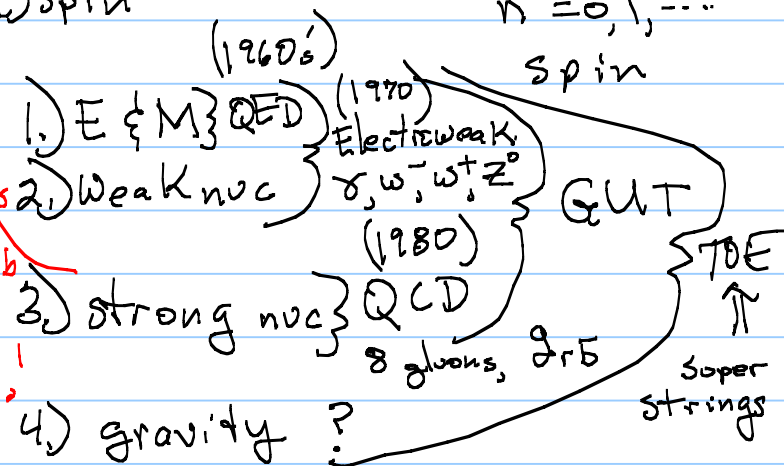
'Particles' = Fermions 'Force' = Bosons

8 generations $n(\frac{1}{2})$ spin

$n = 0, 1, \dots$
spin

neutrinos
M B R
uses
3 each
leptons

2 quarks
Leptons



- 1) e^-
- 2) ν_e
- 3) μ
- 4) ν_μ
- 5) τ
- 6) ν_τ

- 1) u, d
 - 2) u, c, s
 - 3) t, b
- 18 quarks
 u, d, u, c, s, t, b
one of each
color charge!

Super
strings

now 4 ADROWS

Baryons = 3 quark combos

$$p = \begin{pmatrix} u_r & u_g \\ d \end{pmatrix}$$

but colorless!

(Note u_r has \bar{r})

or

$$\Delta^{++} = \begin{pmatrix} u_r & u_g \\ u_b \end{pmatrix} = \text{colorless}$$

* Note: because of this Baryon & Pauli exclusion, clear that addit quantum #, color, was needed

$$\text{Mesons: } = q \bar{q} \text{ combos} \quad \left(\begin{array}{l} \text{need} \\ \text{to} \\ \text{make} \\ \text{colorless} \end{array} \right)$$
$$\pi^+ (u \bar{d}) \quad s^0 (u_g \bar{d}_g)$$

$J/\psi (p\psi c) = (c \bar{c}) = \text{November revolution!}$

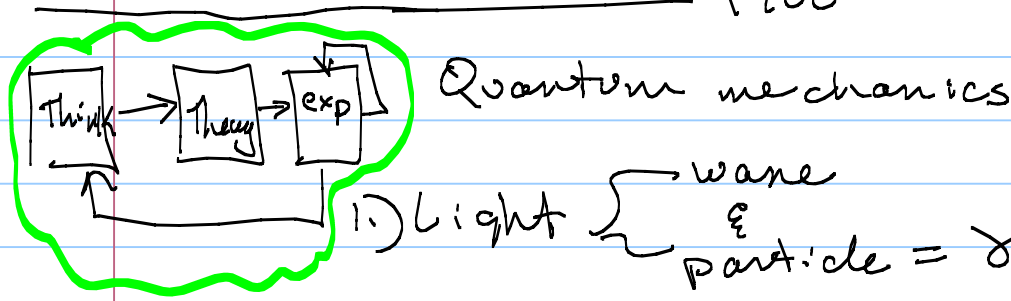
2004: Penta quarks = 5 quark combos

Now

Classical

- 1.) Newton's laws
- 2.) Max's Equations (E.M.)
- 3.) Stat Mech Thermodynamics

1900



1.) Light $\left\{ \begin{array}{l} \text{wave} \\ \epsilon \\ \text{particle} = \gamma \end{array} \right.$

2.) $e^- \left\{ \begin{array}{l} \text{particle} = e^- \\ \text{wave?}, \Psi = \text{wavefunction} \end{array} \right.$

Solve Schrödinger's equation

$$i\hbar \frac{\partial \Psi(x,t)}{\partial t} = \hat{H} \Psi(x,t) \quad (1933 \text{ Nobel Prize})$$

But what is Ψ ?

$\Psi^* \Psi = \text{probability density!}$

Schrödinger - Born

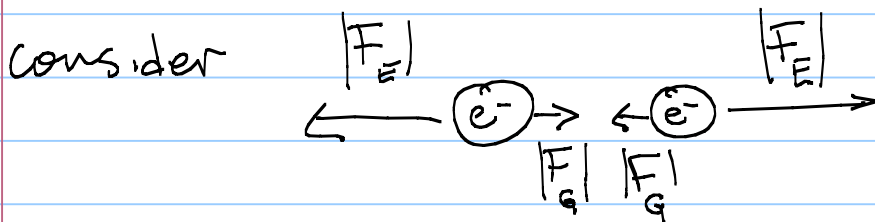
(Max... Olivia's Grand dad)
probabilistic interpretation of Q.M.
1954 Nobel Prize.

Tremendously successful! So apply Q.M. to Forces!

So look @ 4 Forces.....

- 1.) E & M
- 2.) weak nuc
- 3.) strong nuc
- 4.) gravity

Obvious choice... gravity. **BOT!**



$$\frac{|F_g|}{|F_E|} = \frac{G m_1 m_2 / r_{12}^2}{K q_1 q_2 / r_{12}^2} = \frac{G m_1 m_2}{K q_1 q_2}$$

$$m_{e^-} \approx 10^{-31} \text{ kg}, \quad q_{e^-} = 1.6 \times 10^{-19} \text{ C}$$

$$G = 6.67 \times 10^{-11} \frac{\text{Nm}^2}{\text{kg}^2}, \quad K = 10^{10} \frac{\text{kg m}^3}{\text{s}^2 \text{C}^2}$$

$$\left| \frac{F_g}{F_E} \right| \approx \frac{10^{-72}}{10^{-28}} \approx 10^{-44}$$

Relative Strengths & Ranges

Clearly $F_E \gg F_G$

what about

weak nuc, weak & short range $\sim 10^{-15} \text{ m}$

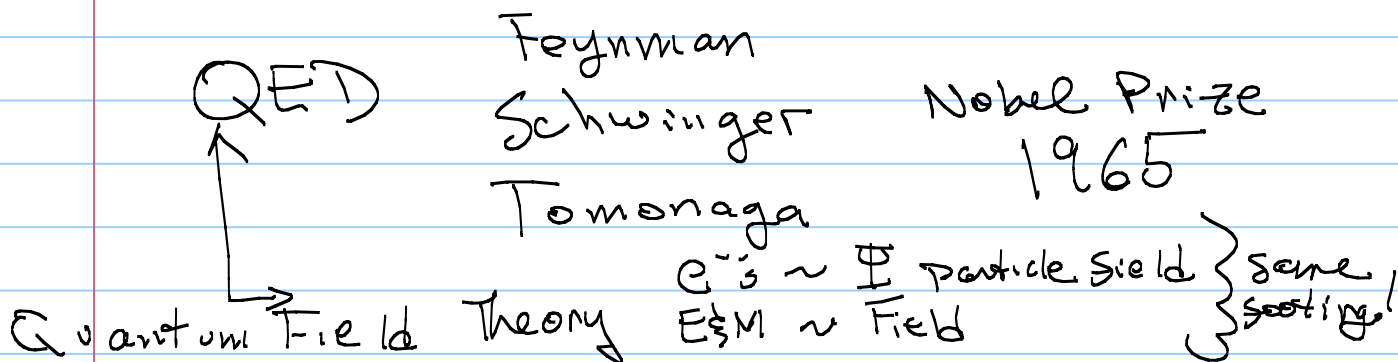
Strong nuc, strong BUT act $\sim 10^{-15} \text{ m}$ of nuc

So, @ the size of

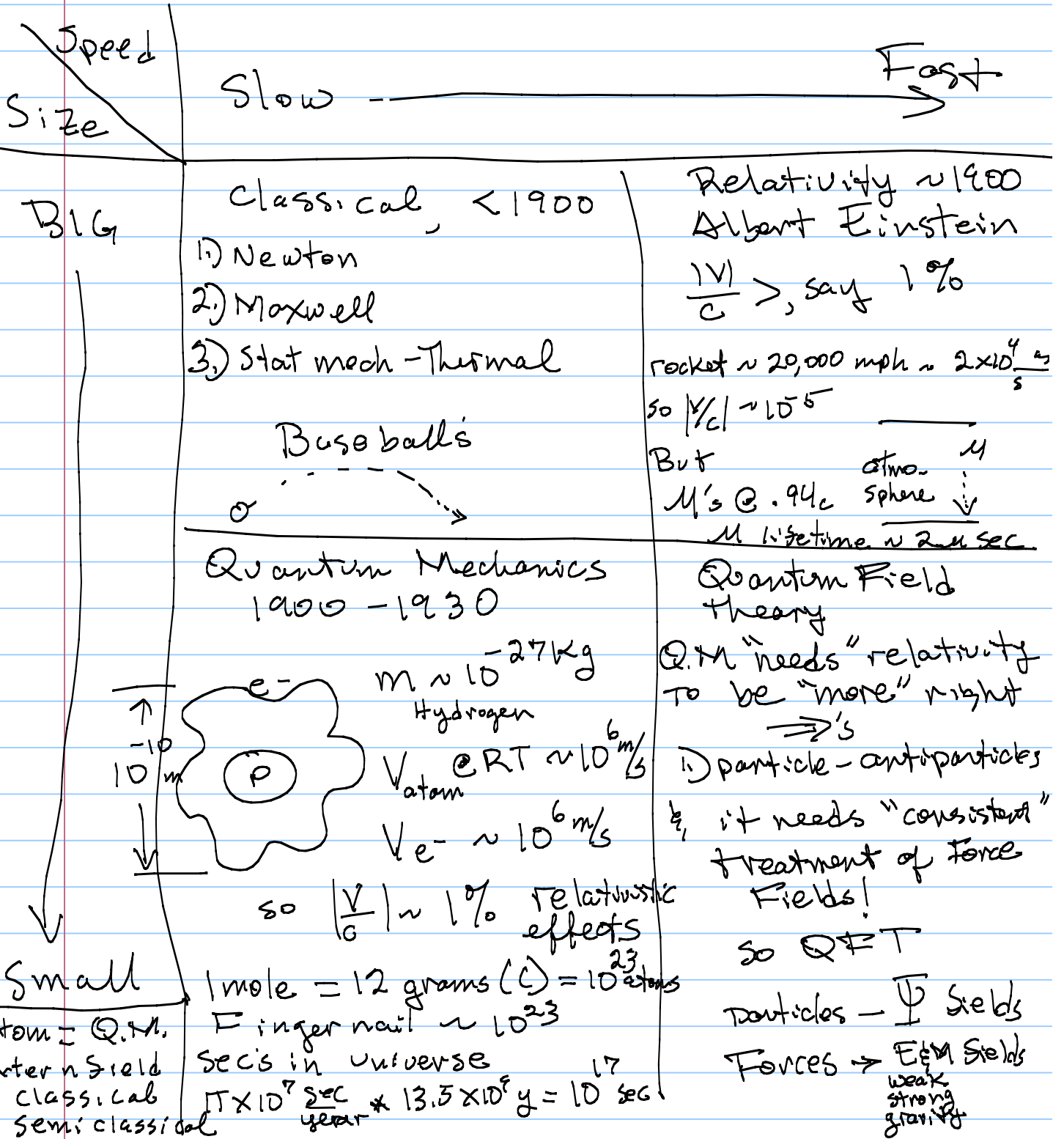
anything $>$ atom

E & M Dominate!

So go about Describing E & M Force w/
Quantum Mechanics 1st



So



Classical, < 1900

- 1) Newton
- 2) Maxwell
- 3) Stat mech - Thermal

Relativity ~ 1900
Albert Einstein

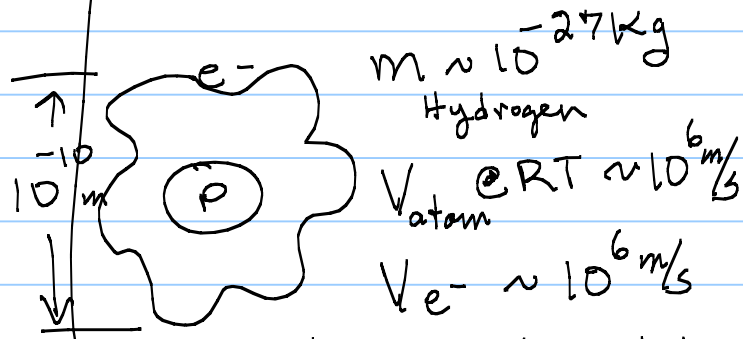
$\frac{|v|}{c} >, \text{ say } 1\%$

rocket ~ 20,000 mph ~ $2 \times 10^4 \frac{m}{s}$
so $|v/c| \sim 10^{-5}$

But M 's @ .94c μ sphere \downarrow
 M lifetime ~ 2 μ sec

Quantum Mechanics
1900 - 1930

Quantum Field
theory



Q.M. "needs" relativity
to be "more" right
 \Rightarrow 's

- 1) particle-antiparticles
- 2) it needs "consistent" treatment of force fields!

so QFT

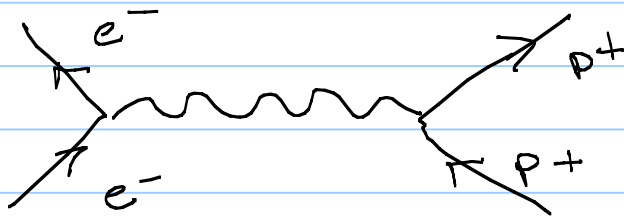
atom = Q.M.
Enter n field = classical
= semiclassical

1 mole = 12 grams (C) = 10^{23} atoms
Finger nail ~ 10^{23}
Sec's in universe 10^{17}
 $\pi \times 10^7 \frac{\text{sec}}{\text{year}} \times 13.5 \times 10^8 \text{ y} = 10 \text{ sec}$

particles - Ψ fields
Forces \rightarrow E&M fields
weak strong gravity

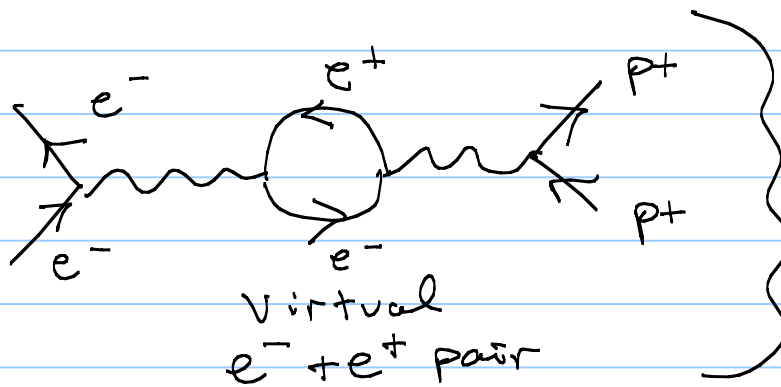
QED most successful, precise theory ever!

H-atom



energy levels
 $\sim 2 \text{ eV} =$
 orange light
 Balmer series
 $n \rightarrow 2$

QED correction



Lamb Shift
 $\sim \frac{1}{10^6} \text{ eV}$

So experiment:

	2.0000003 ←	} wow!
Theory (no Lamb)	2.000000	
Theory (w/ Lamb)	2.0000003 ←	

Great. But deeper? From nothing, can you show that you, the universe, needs to have

Ψ = particle fields

⚡

1	$E \frac{1}{2} M$
2	w nuc
3	S nuc
4	gravity

Force fields

?

Almost!

Gauge Theories

Idea: Invariance \leftrightarrow Symmetries necessitate the Force fields & maybe even the particle fields

How? AJP, 1989, R. Mills "Gauge Theory"

Based on Noether's Theorem

↳ Emmy ... (mathematician)
Student of Hilbert, she had hell of time getting prof job in Germany, forced to flee to US in 1933 and died 2 years later

Brief Sketch of gauge theory

Symmetry \leftrightarrow Conservation; when you "enforce" this symmetry from global to local, you have to add a gauge field so it to work (ie consistent mathemat)

ex: Relativity:

Invariance of "Frames":

To keep the physical laws of motion
[conserv of \vec{p} ; cons of \vec{L} ; cons of E
(invariance of x_0) (invariance of t_0) (invariance of t_0)]

invariant to all frames, inertial & non inertial (accelerating)

need to include GRAVITY & IS you make it Lorentz Invariant local,

Gravity = field that travels @ c

More abstract ...

in Q.M. $\Psi \rightarrow e^{i\theta} \Psi =$
 \uparrow
arbitrary phase

Unitary
Translation!
from Q.M.

Then Schrödinger-Born prob interpret

$$\Psi^* \Psi = e^{-i\theta} \Psi^* e^{i\theta} \Psi = \Psi^* \Psi$$

So Q.M. wavefunction is invariant to $e^{i\theta}$

Is $e^{i\theta(\vec{x}, t)}$ = global, everywhere
sometimes

preserves the "invariance"
even w/ derivatives $\Psi^* \left(\frac{d}{dx}\right) \Psi$

But is demand it to be local

$$\Psi = e^{i\theta(\vec{x}, t)}$$

Then

$$\Psi^* \left(\frac{d}{dx}\right) \Psi \neq e^{-i\theta(x,t)} \Psi^* \left(\frac{d}{dx}\right) e^{i\theta(x,t)} \Psi$$

So loose invariance! or "local" symmetry

Eisberg & Resnick does it well...

time dep Schrö

$$i\hbar \frac{\partial \Psi}{\partial t} = \hat{H} \Psi$$

is transform Ψ globally $\Psi' \rightarrow \bar{\Psi} e^{i\sigma}$

if $\sigma \neq S(x,t)$ then this transformation leads to Schrö invariant.

But... is local $\Psi \rightarrow e^{i\sigma(x,t)} \Psi$

Then

$$i\hbar \frac{\partial}{\partial t} \left(e^{i\sigma(x,t)} \Psi \right) = \left(-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} + V(x) \right) e^{i\sigma(x,t)} \Psi$$

The temporal & spacial derivatives destroy the invariance of Schrö to local transformations!

Free particle

However,

$$\text{is } \Psi \rightarrow (e^{i\theta(\vec{x},t)} \Psi + \vec{F})$$

Then $\Psi \rightarrow \left(\frac{d}{dx}\right) \Psi$ is (completely invariant)
Locally! •

is you investigate what \vec{F}
needs
to be to make Ψ locally invariant,

SURPRISE exactly = E&M field
described by
Max's
Equations! •

where are we?

Need

1.) Invariance of phase \Leftrightarrow cons of charge

γ = photons
of $E\&M$ Force
(QED)

2.) " " Isospin

Need

γ
 Z^0
 W^-
 W^+

} mass
less

problem Not
renormalizable?
(crazy, infinity, answers)

But, I add another field

Higgs field that give mass to
 Z^0 by spontaneous
 W^- broken symmetry
 W^+

The entire theory becomes invariant
AND renormalizable (Gerard 't Hooft
in 1971)

Nobel Prize in 2012?

Wow!

This is electroweak!

γ = photon = $E\&M$

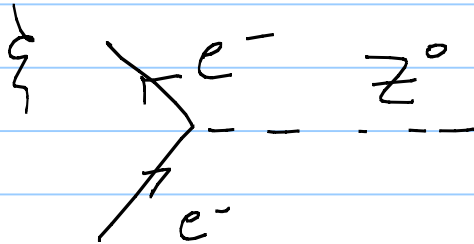
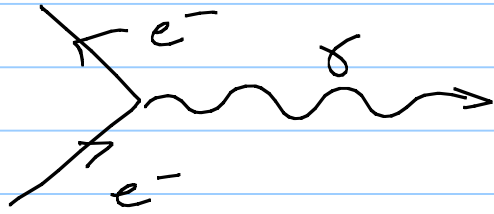
CERN
in the
1980's

Z^0
 W^-, W^+

} weak nuc, massive, short range
neutral & charge currents

see

See



4.) "Color" charge conservation, invariance
quarks can come in (r, g, b)

\Rightarrow is need "Gluon" $g_{\bar{a}b}$ field

8 of them

which explains { all known mesons, π, ρ, ω, \dots
& Baryons, p, n, \dots

1.) asymptotic freedom, why quarks seem "free"
inside of p & n

2.) quark confinement ... yet neither see individual
quarks
(ant. screening ^{quarks} cause gluons carry charge
themselves)

seems like should get the

Quantum Field for gravity! (graviton)

Then The Particles?

Super Symmetry \leftrightarrow all the bosons should have
super symmetric
Fermion 'super partner'

Bosons

photon +ino

Fermion

S + electron

Ex: Superstrings!